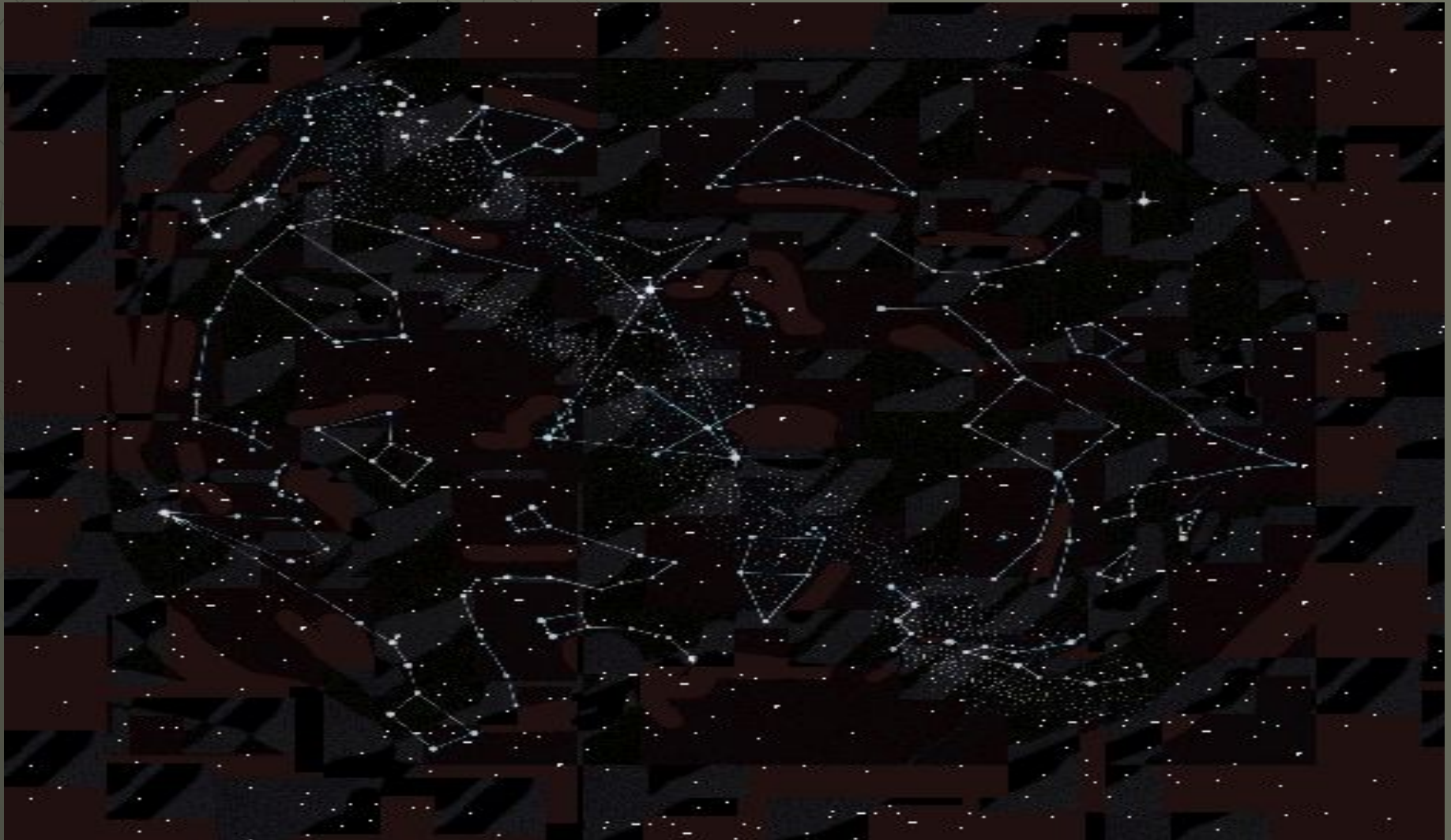


# Life & death of Stars



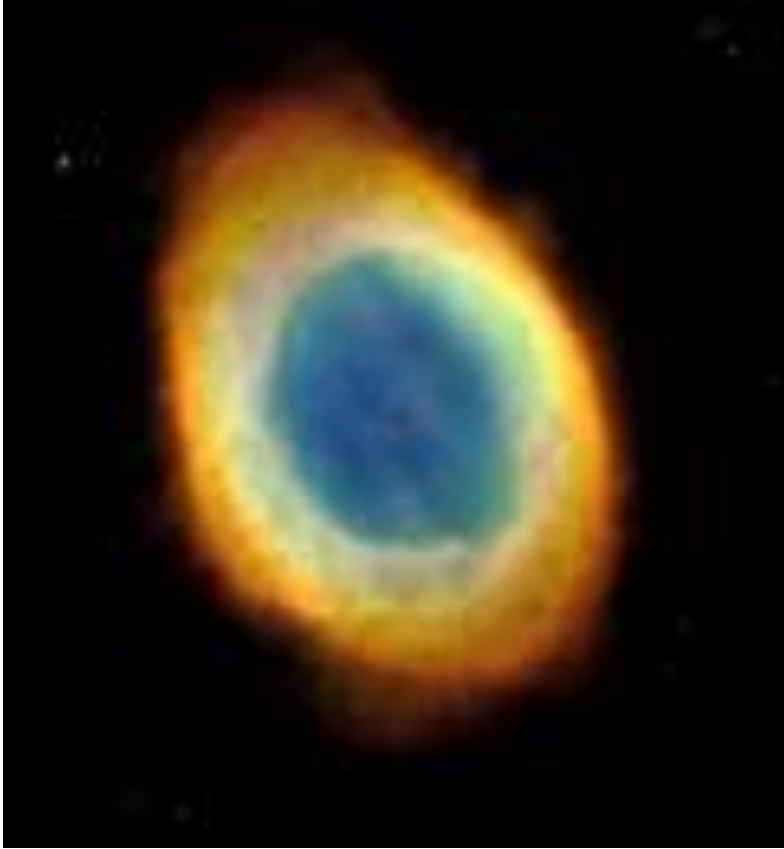
# They're born and then they die....

- ◆ But what gave birth to them?
- ◆ And where do they go when they die?

# *Stages of Stellar Evolution:*



# Nebula (clouds)

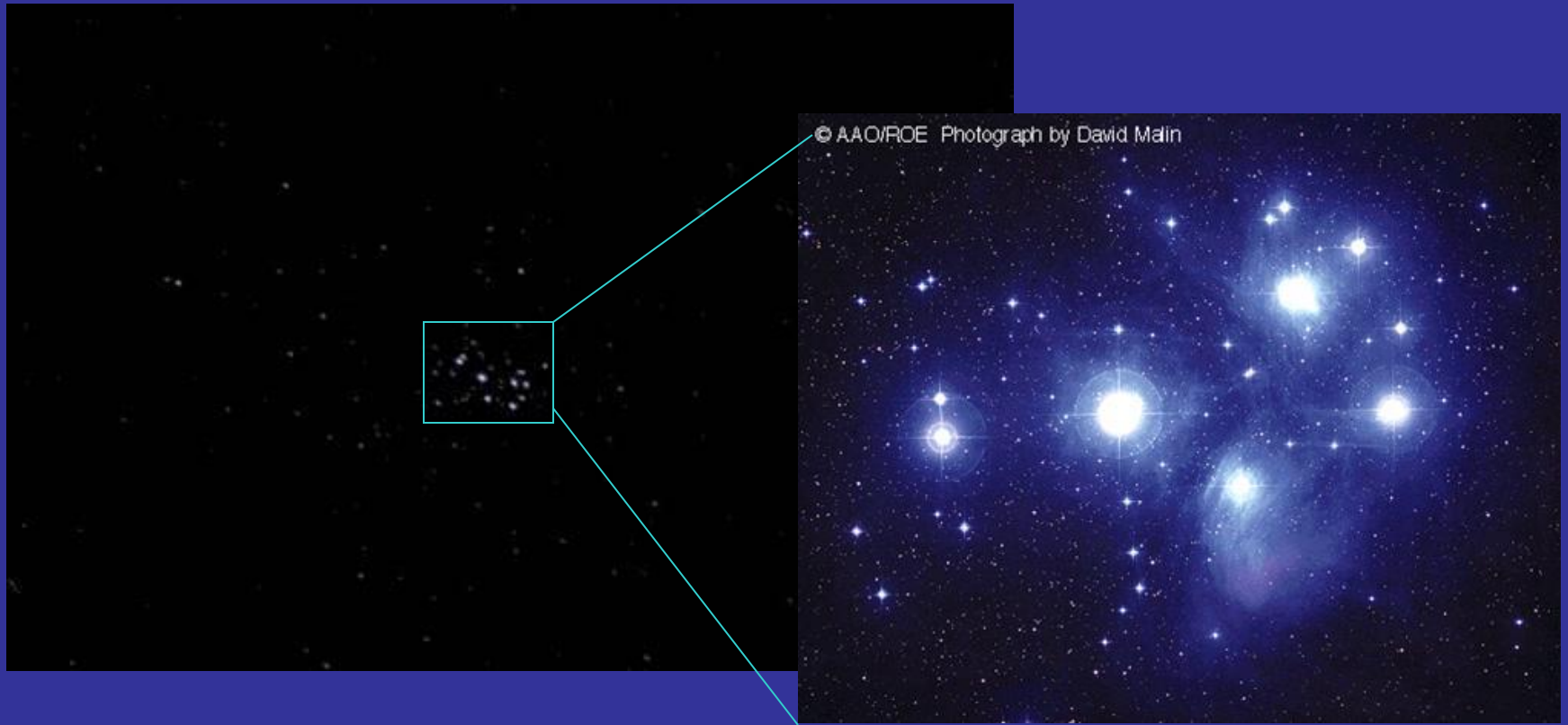


- Nebula - Cloud of swirling dust and gases (mostly hydrogen) and dust --where stars are born

When a disturbance (e.g. supernova shockwave) hits the nebula cloud – a portion of it may begin to contract

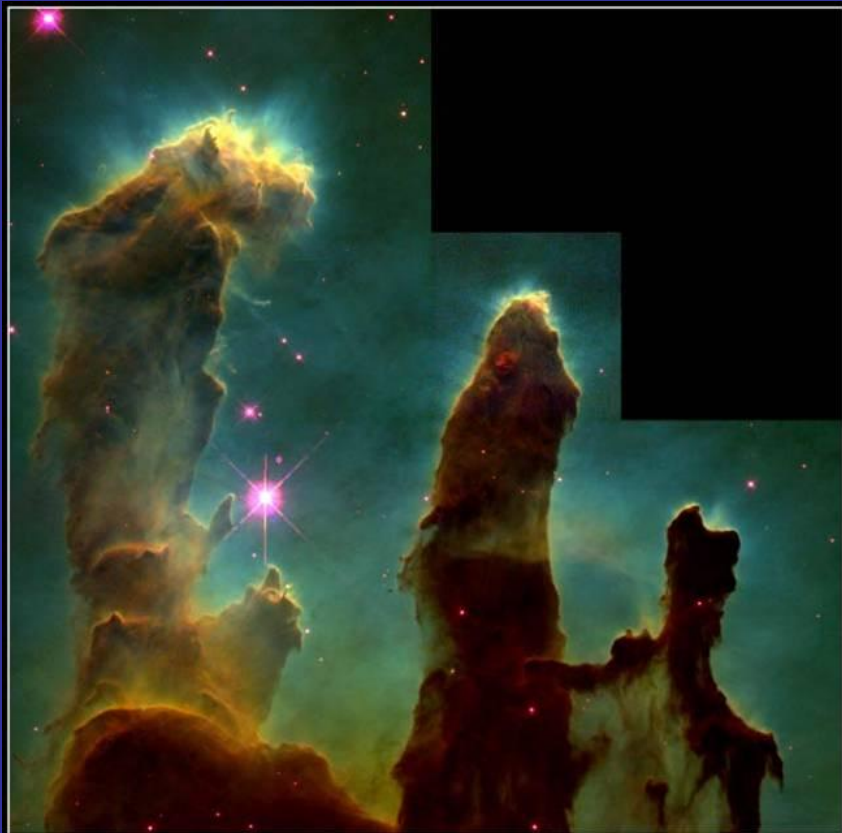
# stellar nurseries

## *Clusters: Pleiades*



# Stellar Nurseries

Optical



**Gaseous Pillars • M16**

HST • WFPC2

PRC95-44a • ST ScI OPO • November 2, 1995  
J. Hester and P. Scowen (AZ State Univ.), NASA

Infrared: 30 Doradus in LMC



**The Tarantula Nebula**

Spitzer Space Telescope • IRAC

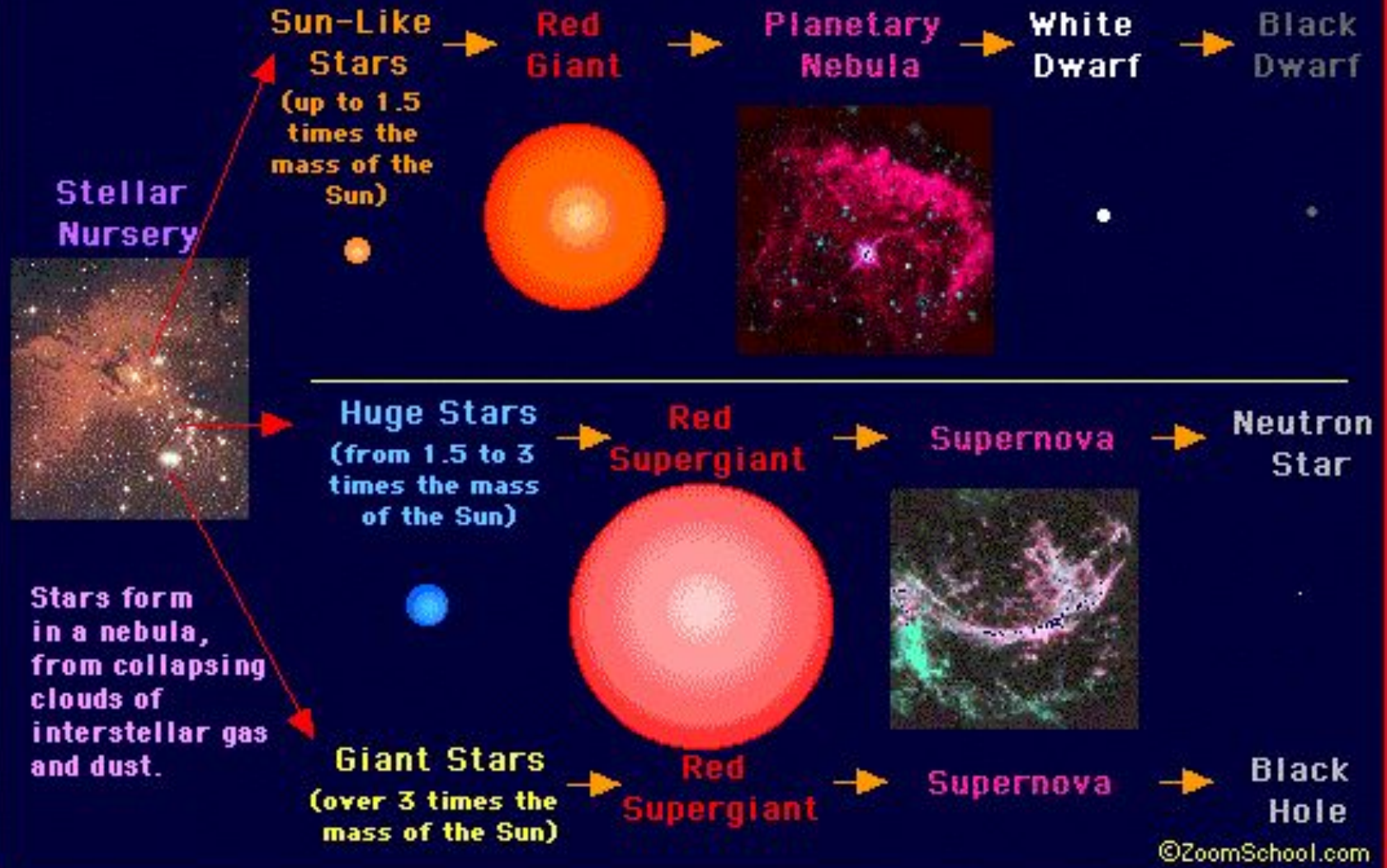
NASA / JPL-Caltech / B. Brandl [Cornell University & University of Leiden]

ssc2004-01a

May 2006 April 2004

Belinda Wilkes

# The Lifecycle of Stars



# PROTOSTAR



- Once a clump of gas and dust begins to contract, gravity becomes stronger  dust and gas collapse  protostar.



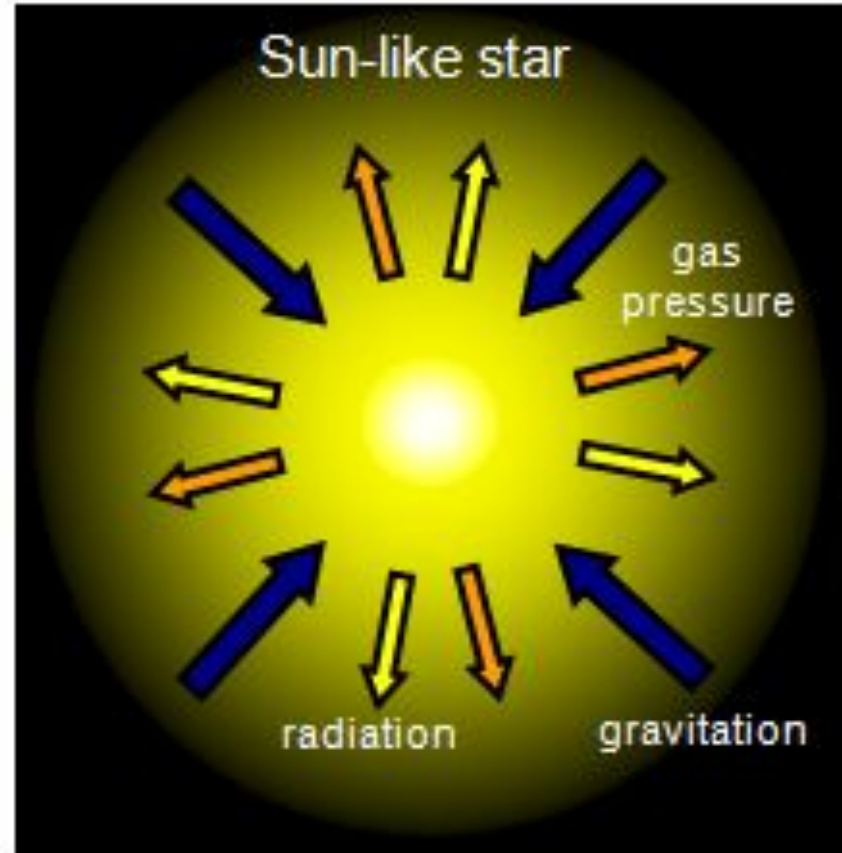
- 
- A vibrant nebula with green and yellow filaments against a dark starry background. The nebula features a prominent yellow and orange filament on the left side, and a large green structure in the center and right. Numerous bright stars are scattered throughout the dark space.
- As it gets collapses and gets smaller the protostar gets hotter

# Main Sequence (Stars)

- WHEN THE PROTOSTAR IS DENSE & HOT ENOUGH IT WILL BEGIN nuclear FUSION which releases large amount of energy
- Nuclear Fusion -- HYDROGEN IS BEING CONVERTED TO HELIUM
- IT IS NOW A MAIN SEQUENCE STAR!

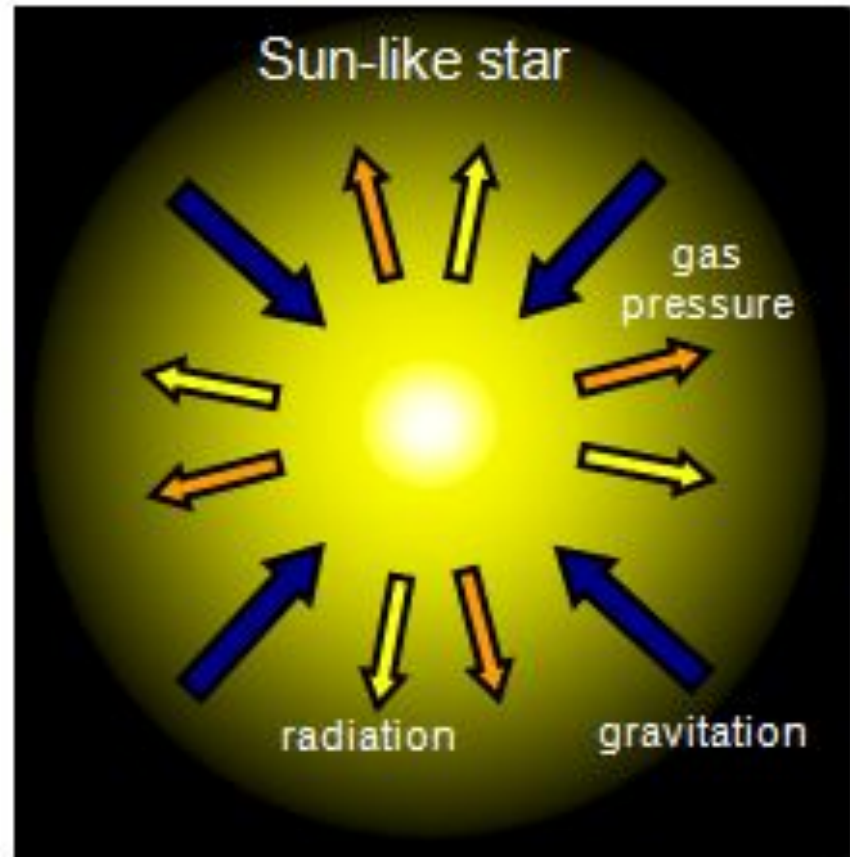
# Forces on a star

- Inward forces
  - Gravitation attraction between particles
  - contraction
- Outward forces
  - Heat and radiation from nuclear fusion
  - expansion of star



If the star is  
stable,

inward  
gravitational  
forces = outward  
forces of gas  
pressure and  
radiation.

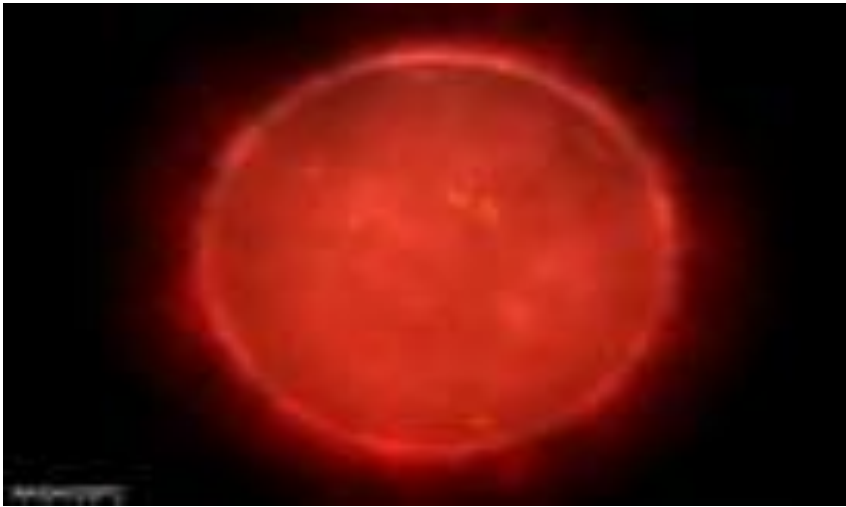


There are three types of Main sequence stars: Low Mass star (red dwarf), Medium Mass (yellow star) or High Mass (blue giant).



# Low Mass MAIN SEQUENCE STARS

called -(RED DWARF) STARS



- A small star that lives billions of years
- Mass less than 4 times the mass of our sun

# Our sun is a Low Mass Star



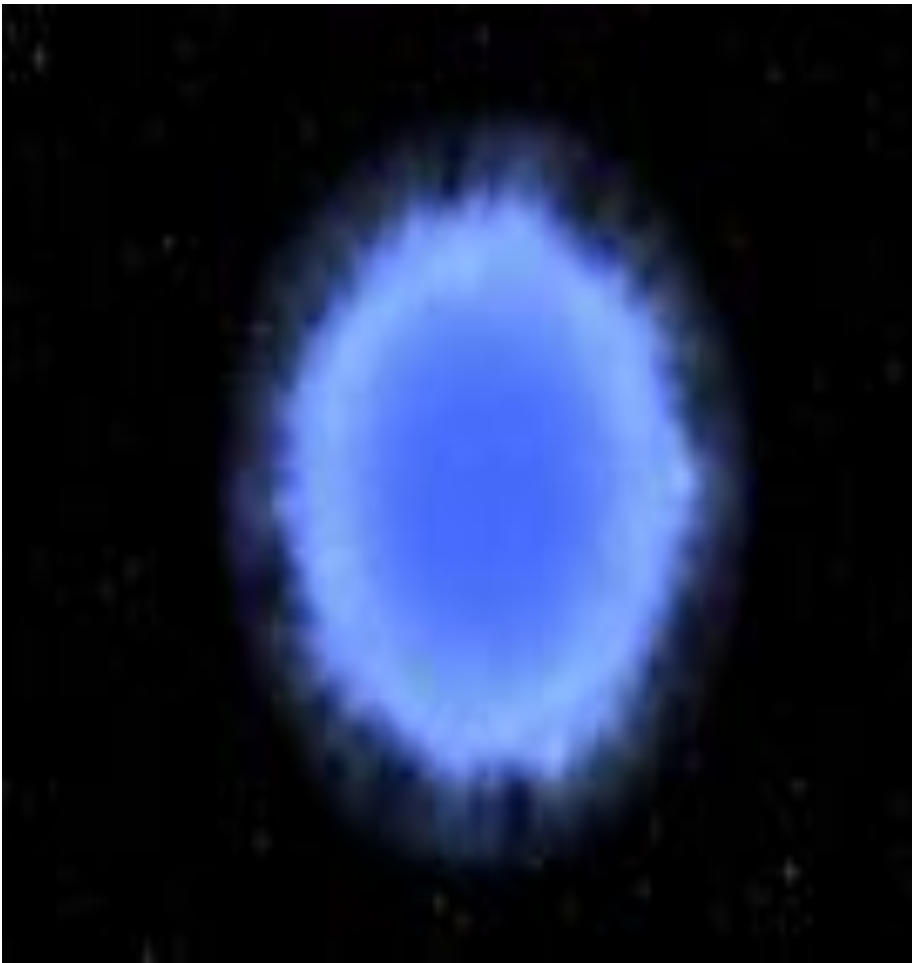
# Medium Mass - MAIN SEQUENCE STAR – called a YELLOW STAR



- A medium mass star
- (4-10 times mass of our sun)
- Shorter lived than a red dwarf but still lives billions of years

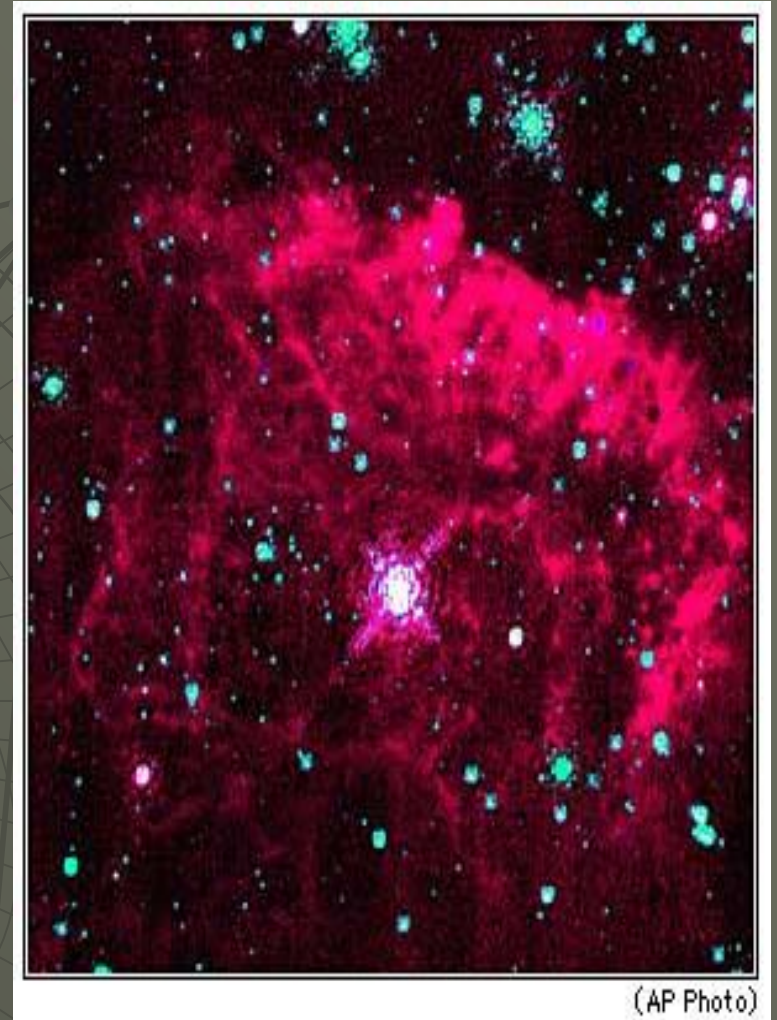


# High Mass (high mass) - MAIN SEQUENCE SIZE BLUE GIANT STAR



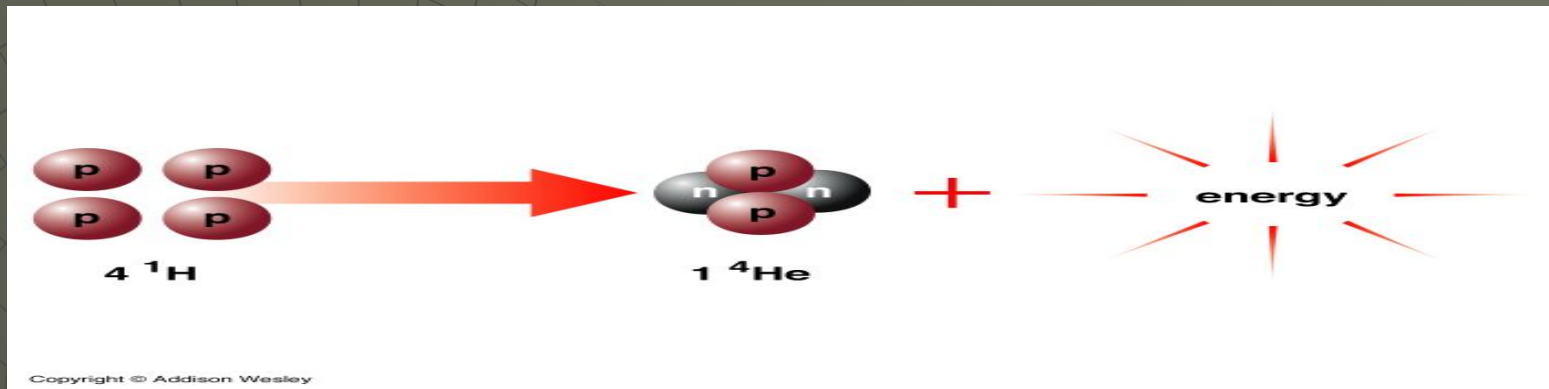
- A large star ( $>10x$  mass of our sun)
- It has the shortest life span
- Living only millions of years.....

# Blue (large) Stars



(AP Photo)

# Transition between Main Sequence and Red Giant/Superbiant



- ◆ As the Hydrogen  $\square$  Helium nuclear fusion reaction runs out of hydrogen  $\square$  gravity (inwards force)  $>$  fusion heat – (outwards force)  $\square$
- ◆ Star starts to shrink  $\square$

# Stars collapse □ then expands

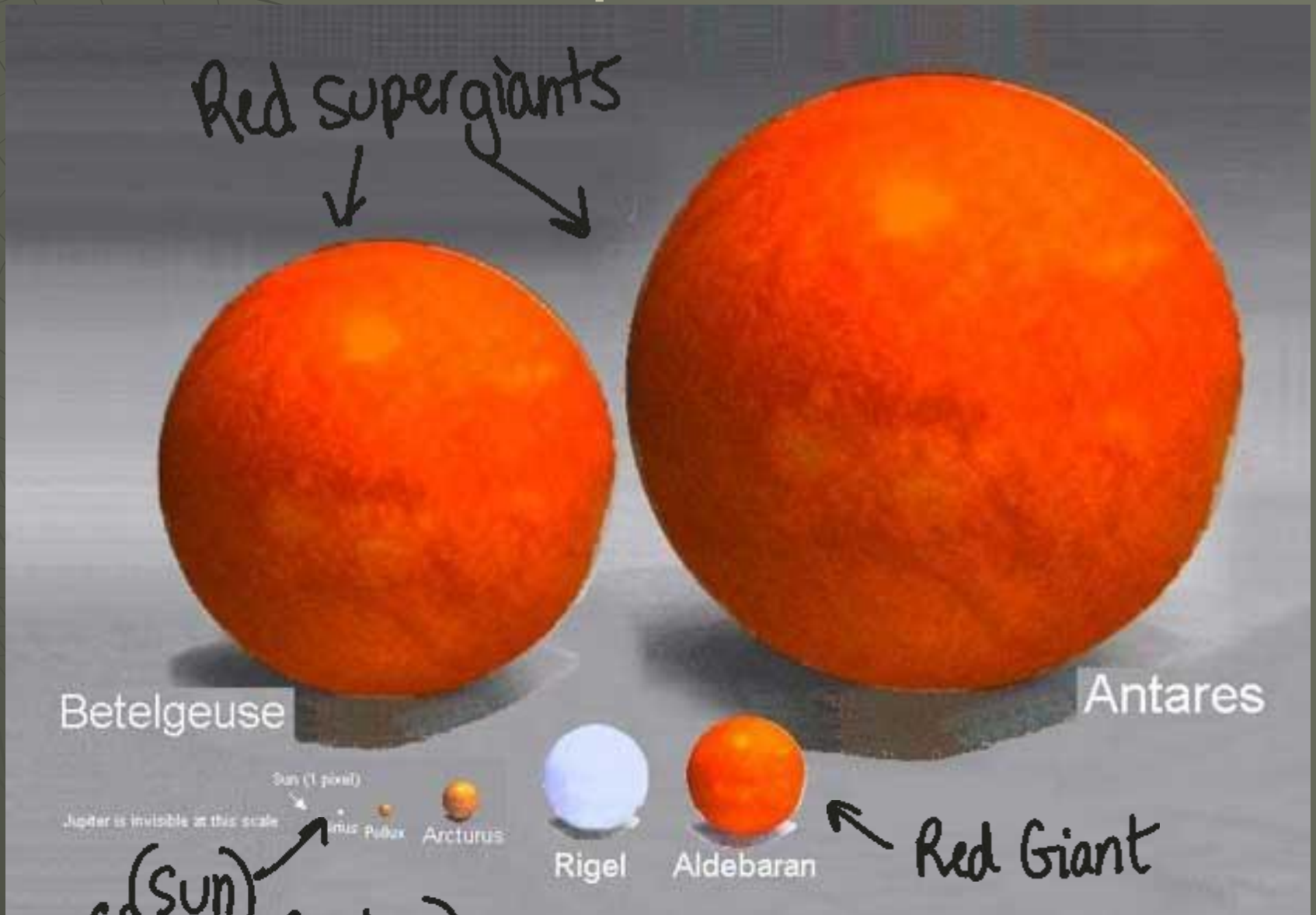
- ◆ Shrinking star □ an increase in pressure and temperature □ restarts nuclear fusion with larger molecules. an increase in radiation and luminosity and the star expands □ red giants

# Red Giant Stars

Low Mass stars (Red dwarf)  $\square$  becomes a red giant



# Lets compare sizes...



(Sun)  
(for comparison)

# Red Giant



Size of the star

or

Orbit of Earth



Orbit of Jupiter

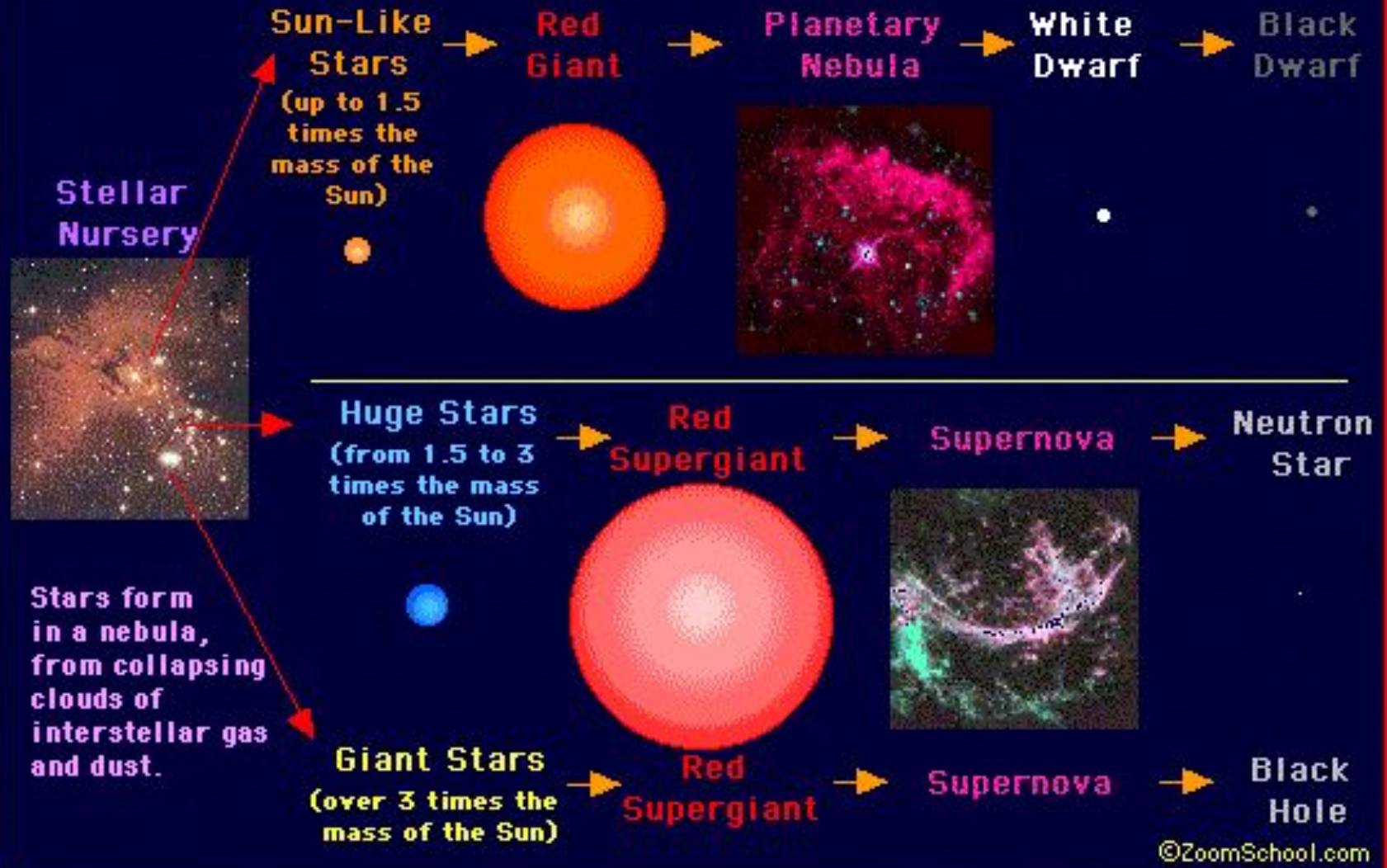
# RED GIANTS

- Star turns red – due to cooler temperature.
- Star increases in luminosity (brightness) due to increased size
- Hydrogen has been used up – so Fusion – involves matter fusing into bigger elements
- Helium converts to carbon & oxygen
- Causing star to expand





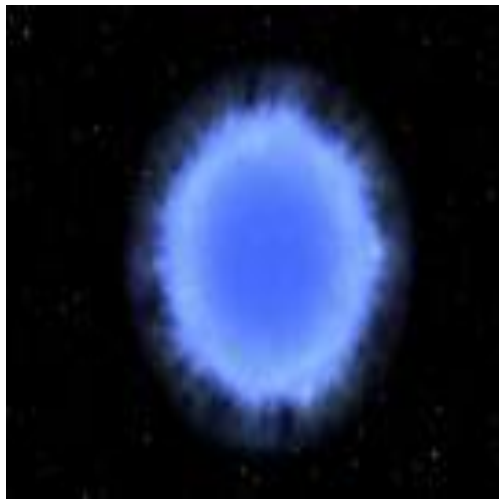
# The Lifecycle of Stars



# RED SUPERGIANT STARS:

Medium & Large Mass stars

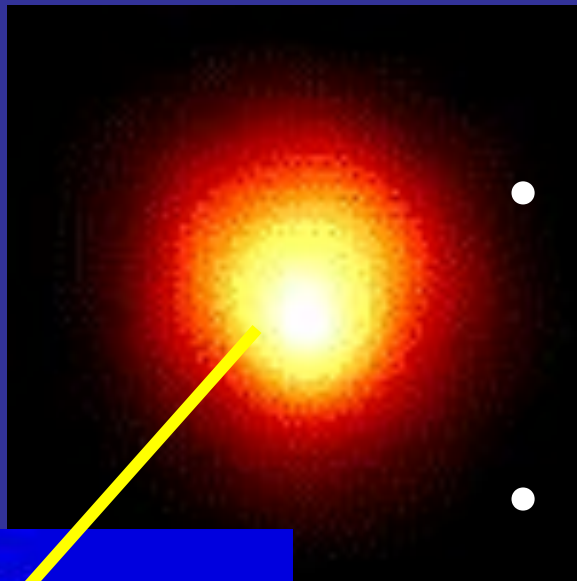
□ becomes a red super giant



# RED SUPER GIANT



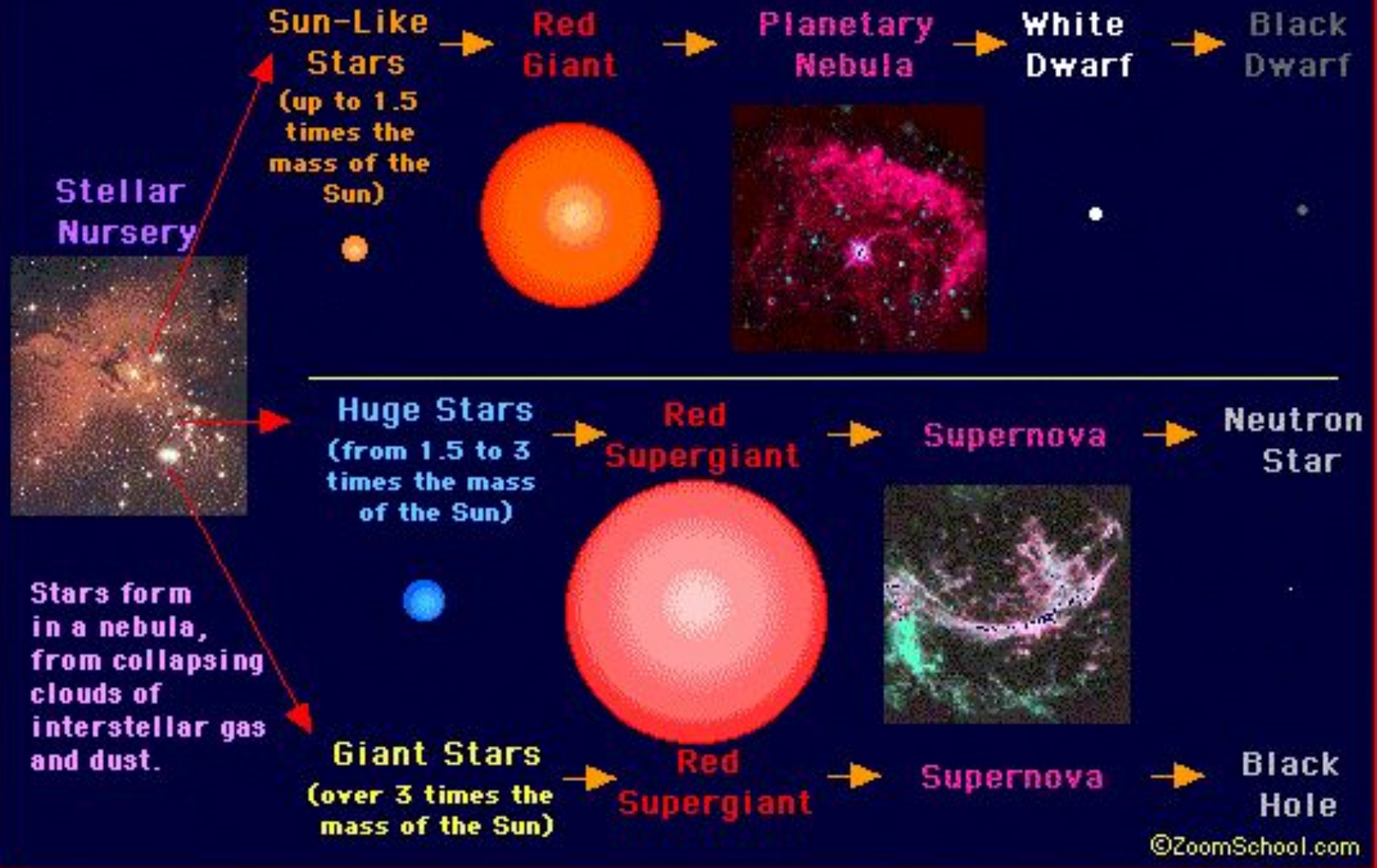
- Loses energy ☐ cooler ☐ red colour
- Larger size ☐ increased luminosity
- Nuclear Fusion He ☐ Carbon ☐ oxygen
- It expands engulfing nearby objects
- Fusion ☐ heat ☐ further expansion



- Example: Betelgeuse is a red supergiant in the constellation of Orion.
- It is over 600 million miles in diameter (1,000 times bigger than the Sun but cooler).
- If Betelgeuse were at the centre of our Solar System, it would extend beyond the orbit of Jupiter.
- It is 520 light-years from Earth.



# The Lifecycle of Stars

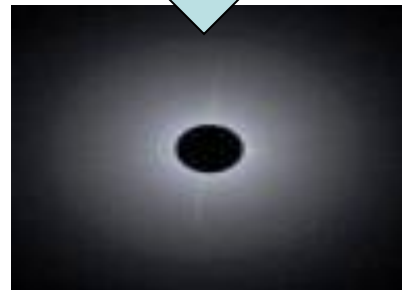
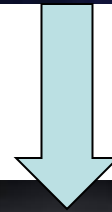


# Star DEATH

- ◆ When the red giant star is no longer undergoing fusion, it begins to collapse – forces of gravity  $\gg$  outward forces .
- ◆ What happens depends on the original mass of the star.

# 1. Low Mass Stars □

Red giants □ a planetary nebula □ a  
white dwarf □ a black dwarf





## *Planetary Nebula*

a kind of emission nebula consisting of an expanding, glowing shell of ionized gas ejected from old red giant stars late in their lives.



# A. WHITE DWARF



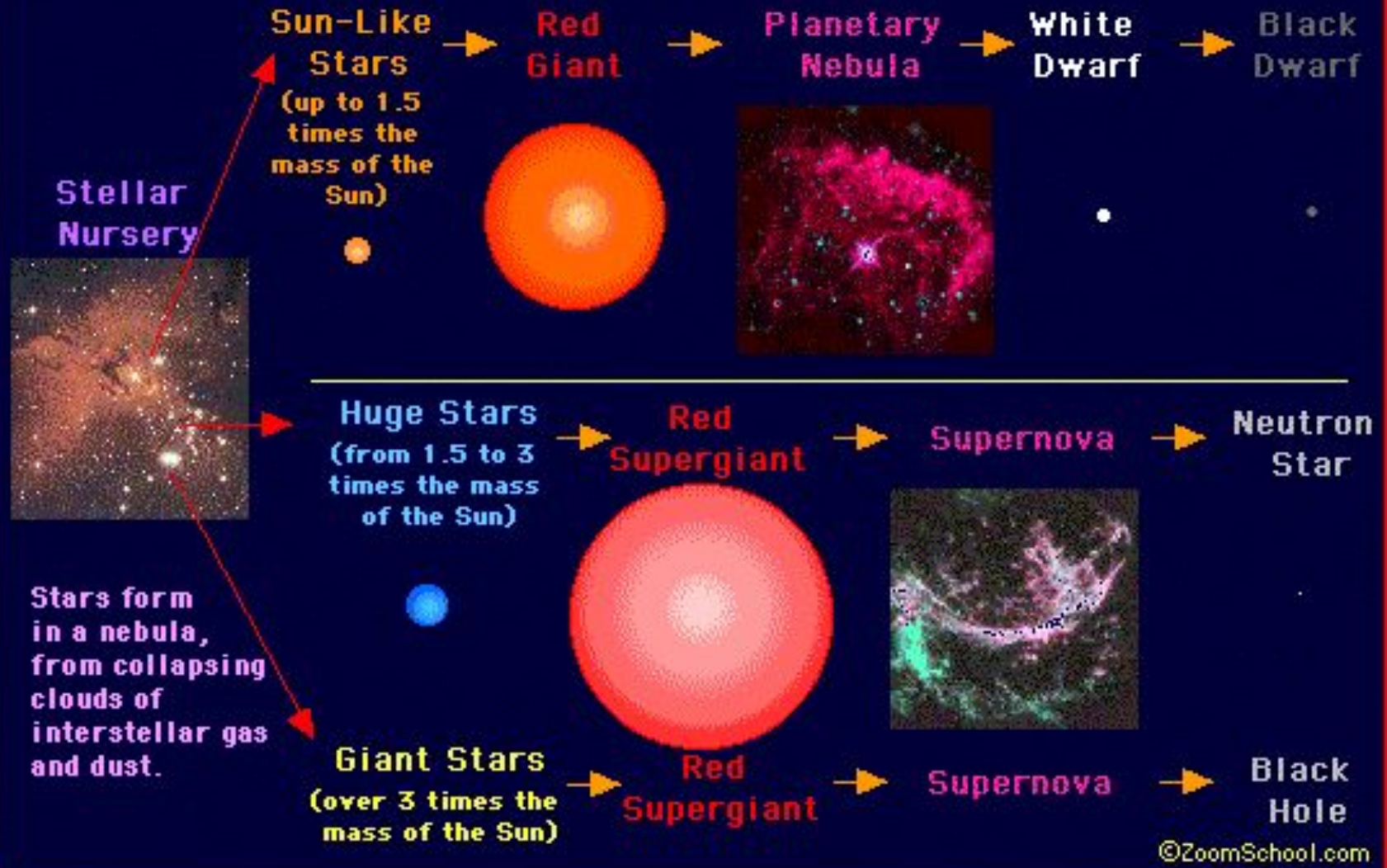
- It cools and contracts
- Helium runs out & core contracts,
- outer layers leave
- produces this small dense star

## B. BLACK DWARF



- Once the White Dwarf has lost all of its heat to space, it no longer radiates heat or light and is cold and dark.
- It becomes a big “charcoal piece” in space

# The Lifecycle of Stars



## 2. Medium Mass Stars:

### A. Supernova neutron star

- A **supernova** is the death of a large star. It is a spectacular explosion (appears millions of times brighter).
- Lasts only for a few days. Maybe be visible during the day.

Supernova are rare – once every century in a typical galaxy.

The remnants of a  
supernova in the  
constellation Cassiopeia,  
all that can be seen by  
astronomers.

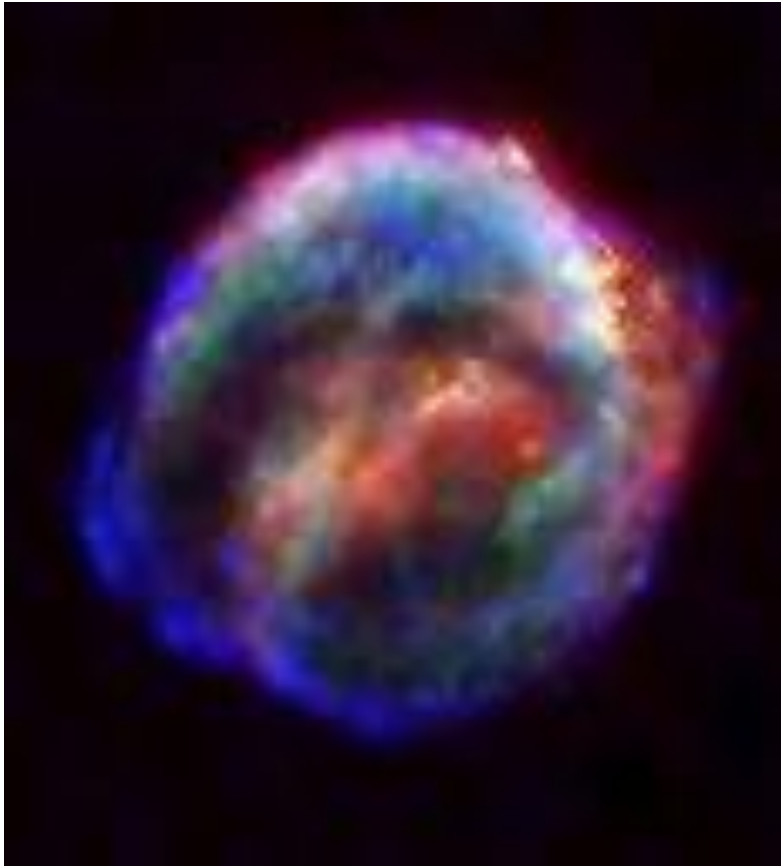
*But the core remains...*



# Supernova

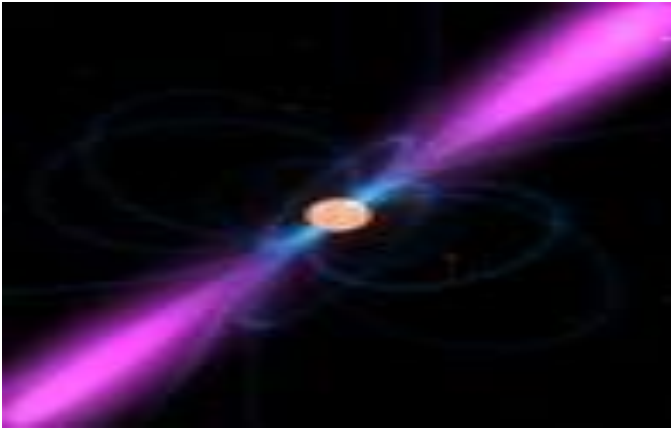
- They emit visible, infrared and X ray radiation.
- Temperatures rise to 10 billion K.

# SUPER NOVA



- Enough energy to cause medium weight elements to fuse, forming heavy elements (up to Uranium in the Periodic Table).

# B. NEUTRON STAR (Pulsar) – (from Medium Mass stars ☺)



- A very dense dead star
- Formed from mass left over after the supernova.



*Calvera, the closest  
neutron star found  
in Ursa Minor*



The gravity is so strong that the electrons and protons combine to form neutrons ☐ Neutron Star

# 3. High Mass Stars

## A. Large Supernova

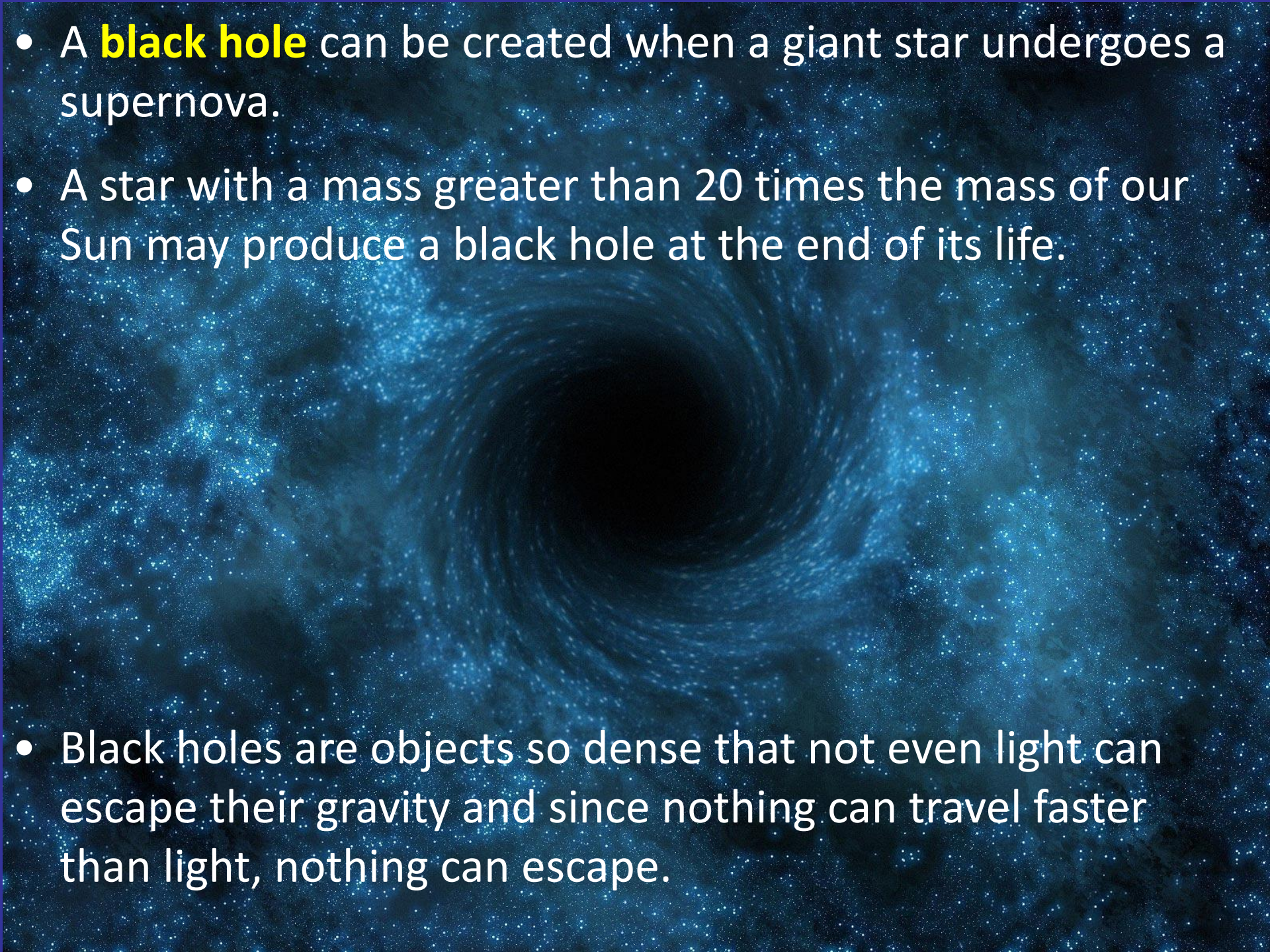


### 3. High Mass stars

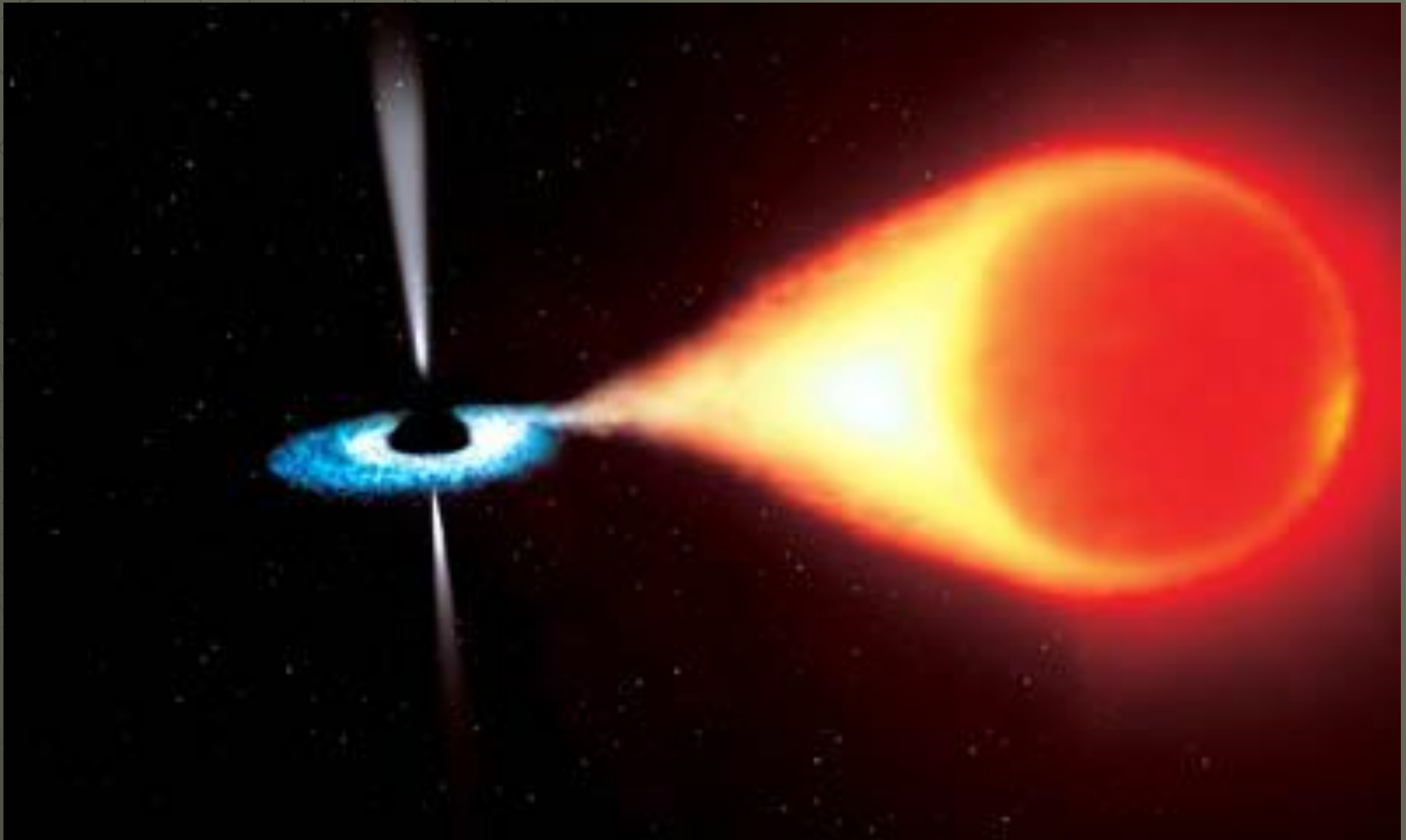
#### Large Supernova

##### B. Black hole

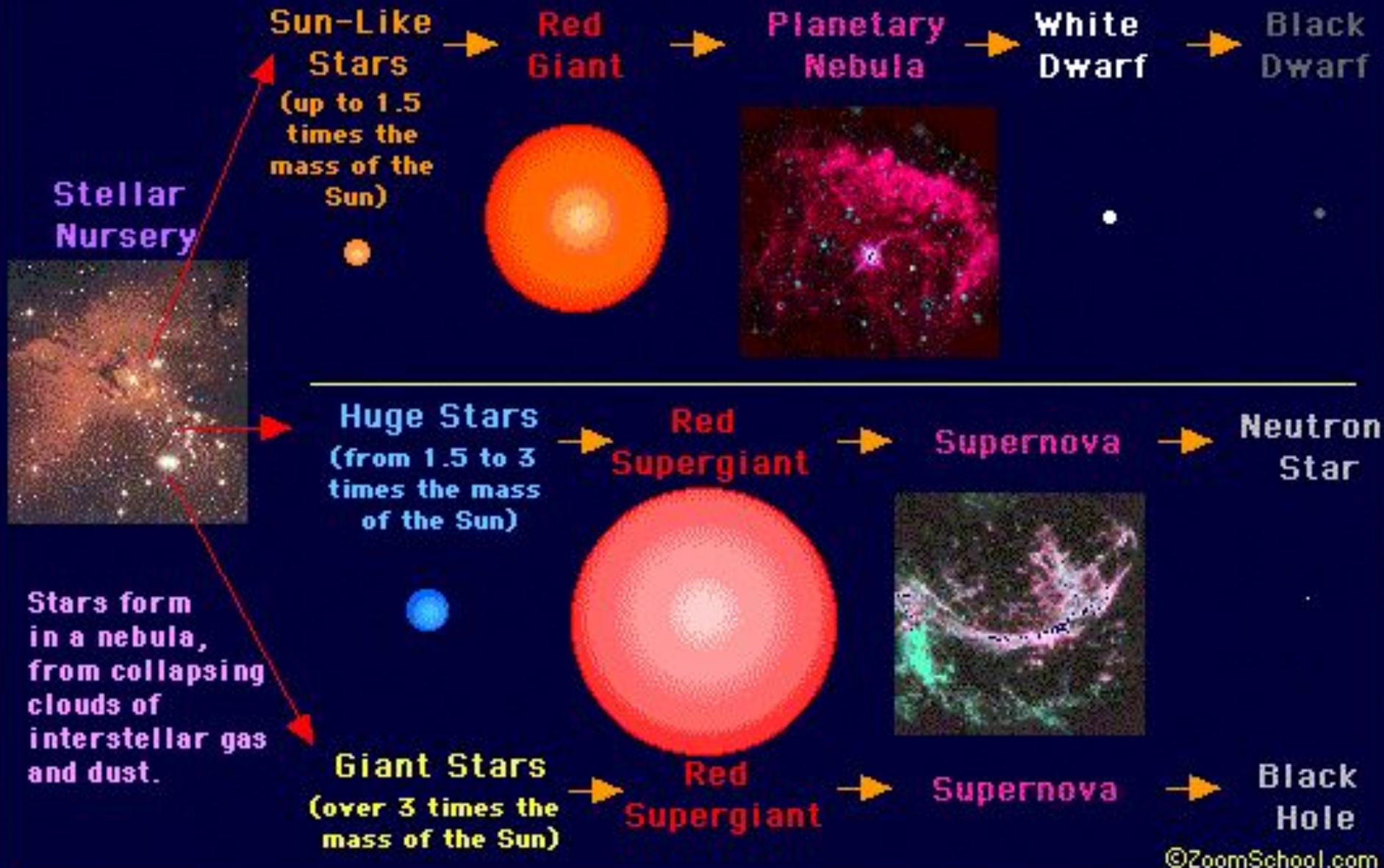
- ◆ If the star is very large, the left over star after the supernova is so dense that it forms a *black hole*
- ◆ Its gravity is so great that light can not escape from it.

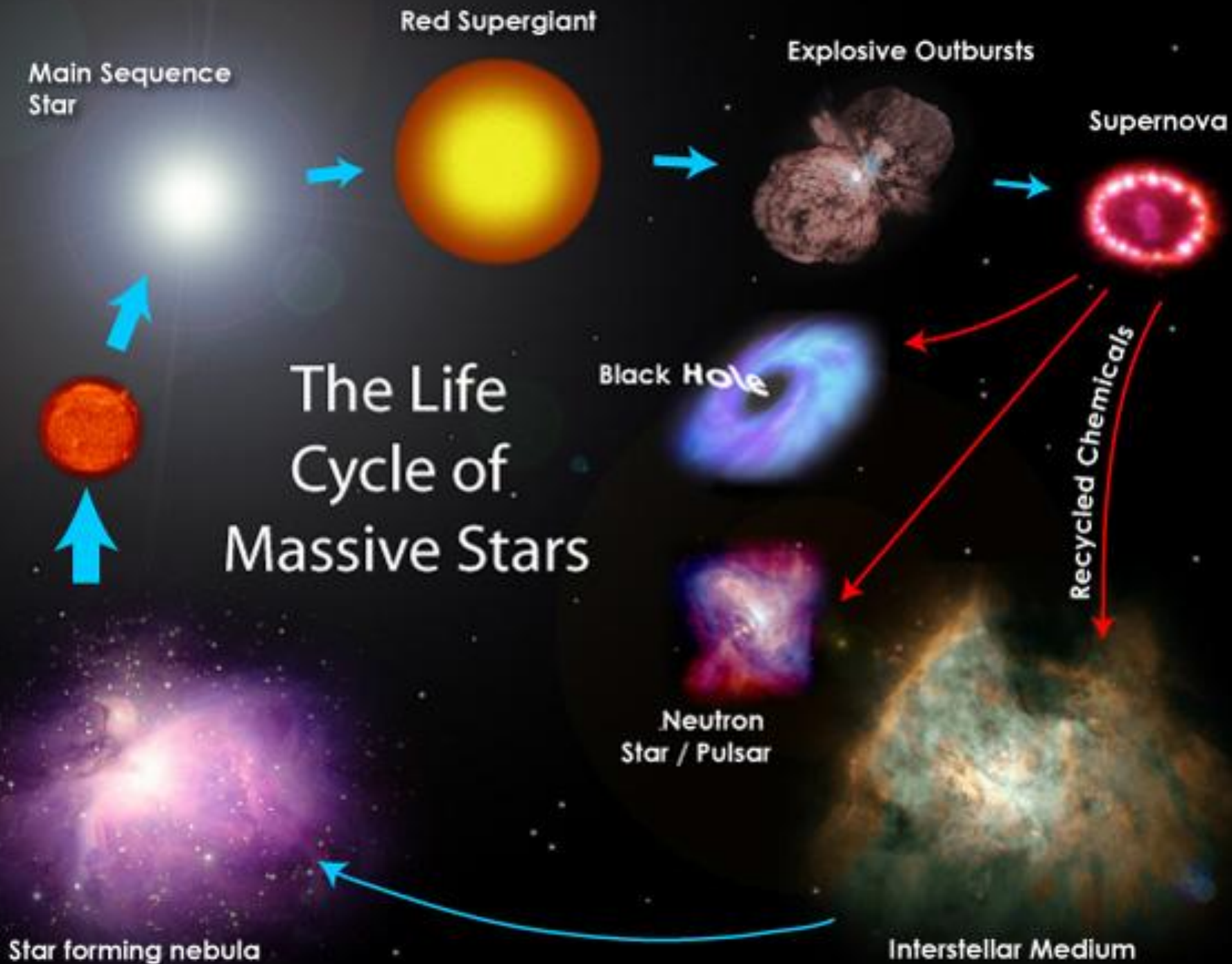
- 
- A swirling black hole in space, surrounded by a field of stars. The black hole is a dark, circular region with a bright, glowing ring of light around it, set against a background of numerous small, bright stars.
- A **black hole** can be created when a giant star undergoes a supernova.
  - A star with a mass greater than 20 times the mass of our Sun may produce a black hole at the end of its life.
  - Black holes are objects so dense that not even light can escape their gravity and since nothing can travel faster than light, nothing can escape.

Black Holes – only High mass stars will end up as black holes



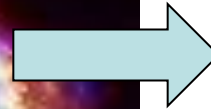
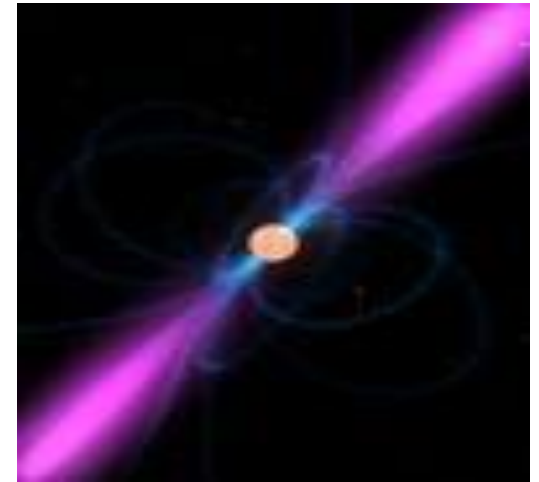
# The Lifecycle of Stars



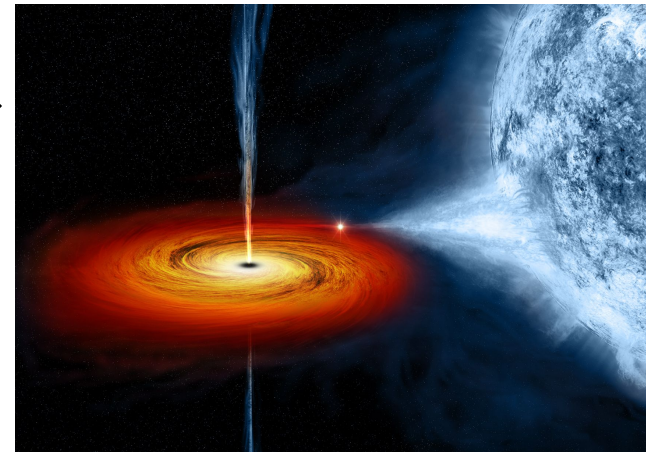


Summary - Red super giant  $\square$  a  
super nova  $\square$  either a neutron star  
or a black hole

Neutron star

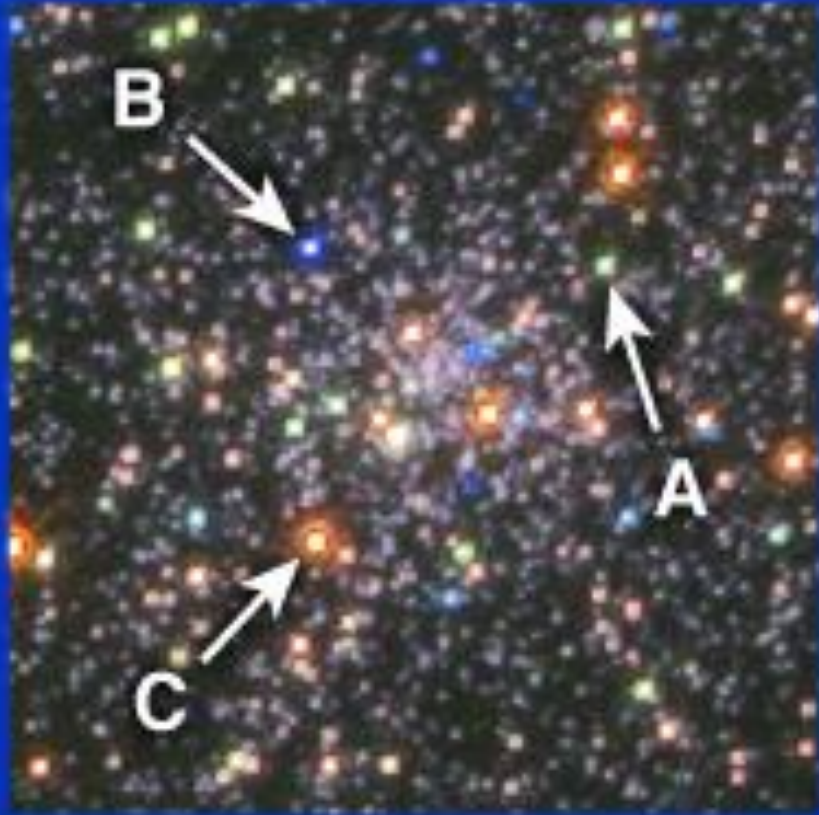


Black hole



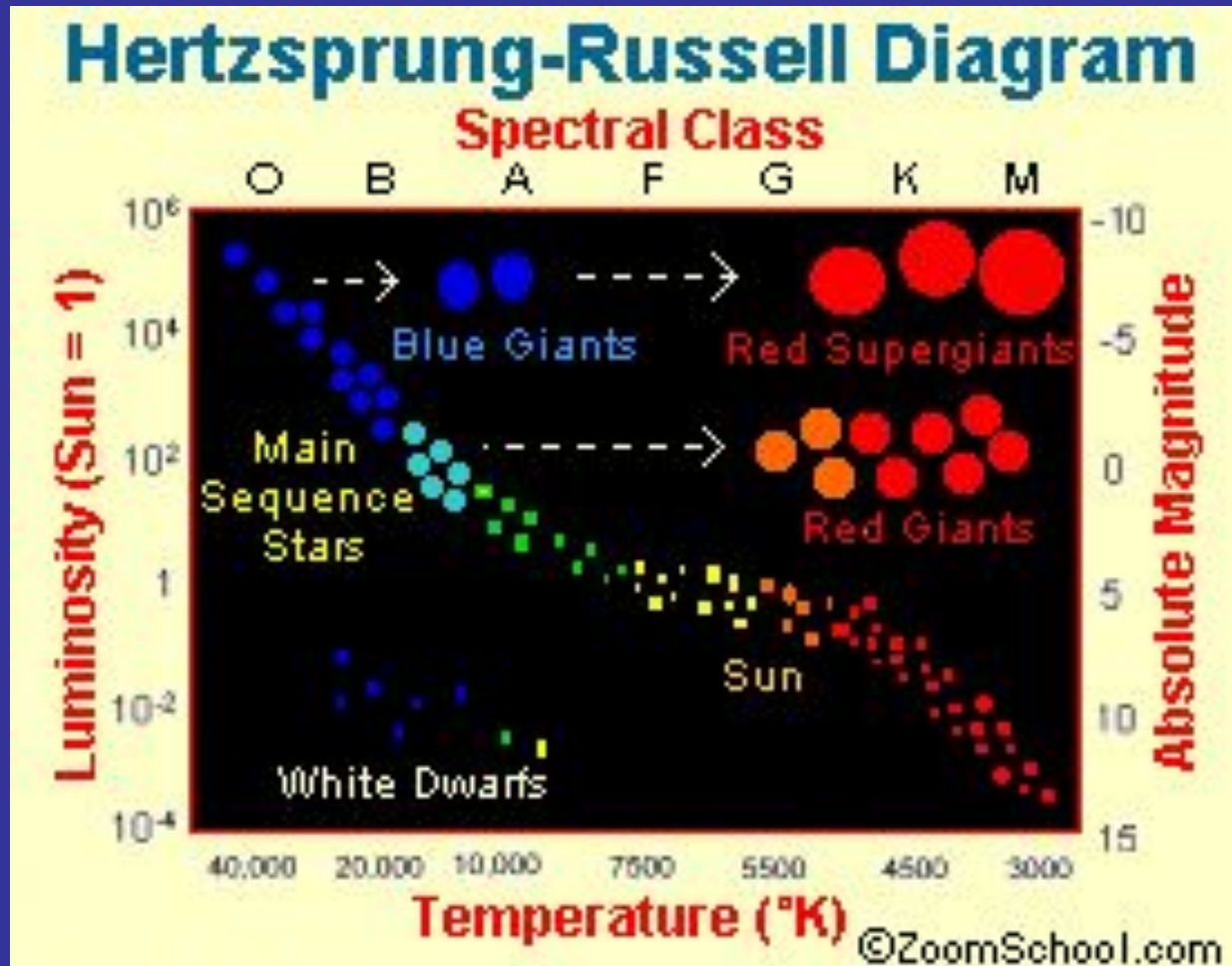


# Stars colour relates to its temperature



- B: blue – hottest
- A: green – warm
- C: red - cool

# Color, Brightness + Count them

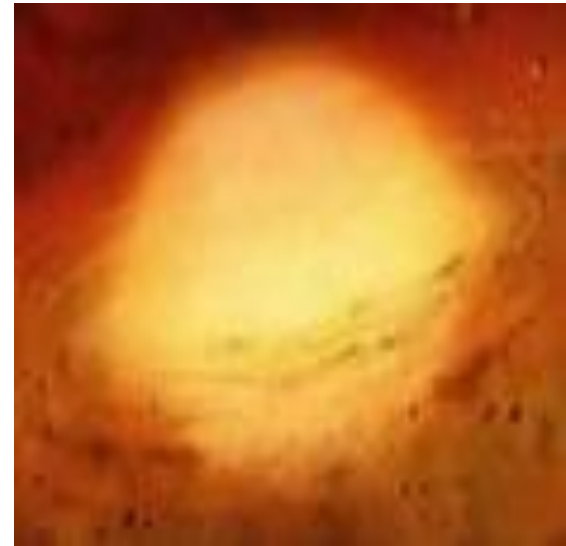
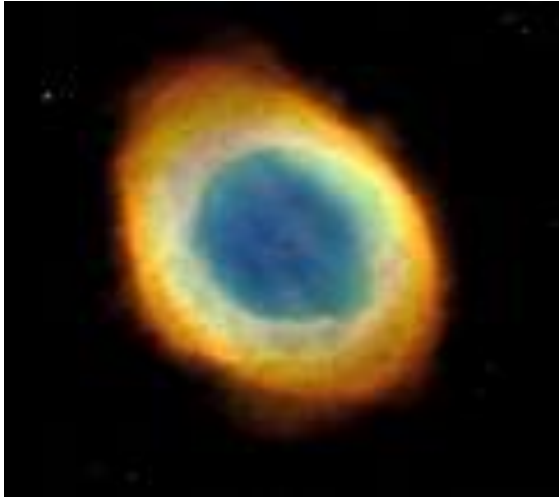




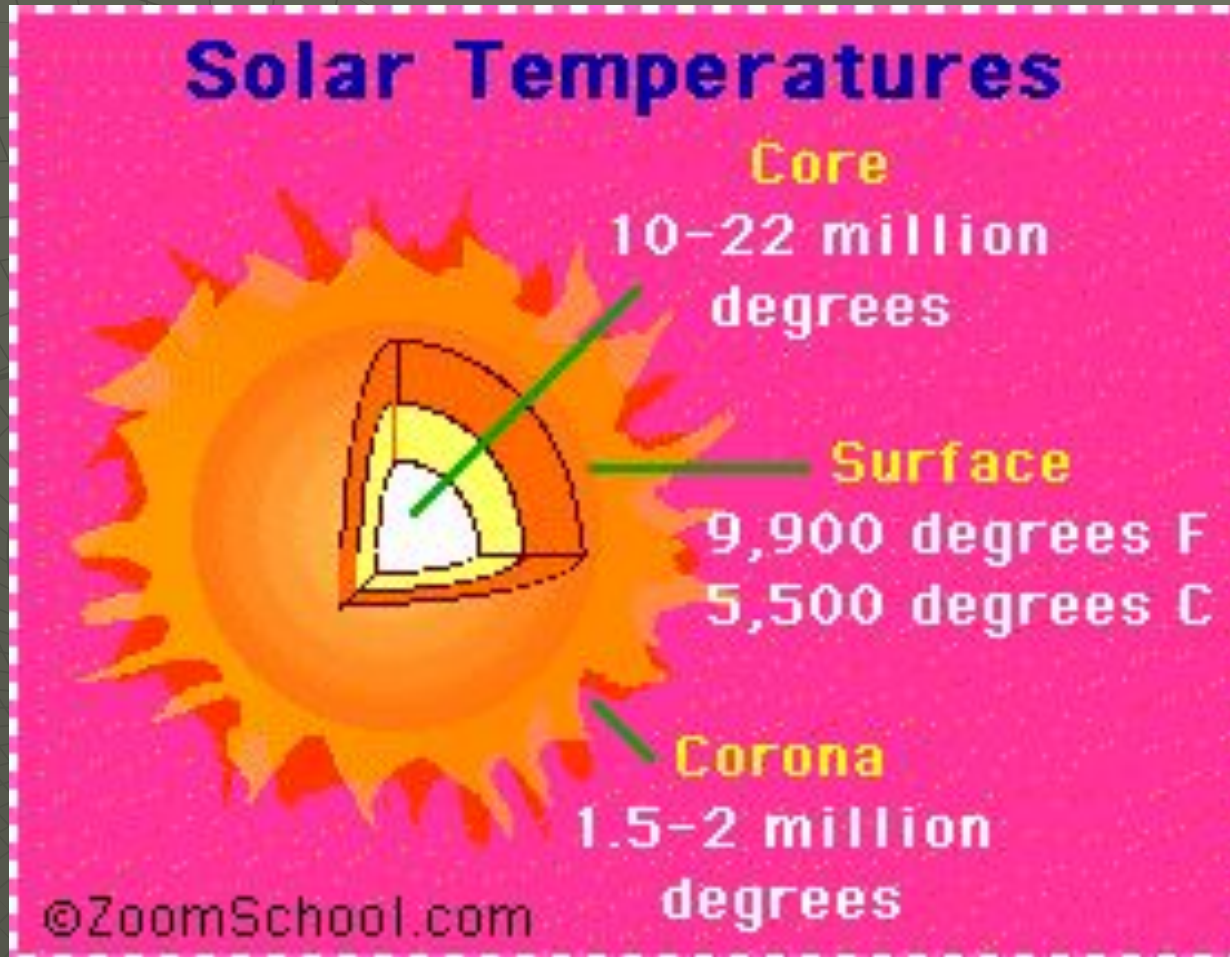
- A planetary nebula occurs at the end of a red giant's life.
- The outer layers of the red giant start to drift off into space.
- This is *The Eskimo Nebula*

# REVIEW OF STARS LIFE CYCLE

## Nebula to proto star



# What makes the Sun Shine?



# How stars burn

- Main Sequence Star – use nuclear fusion -- H as fuel  
□ He
- Giant star – He as fuel □  
O, C, Fe

